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Key Points:

- Antarctic topographic change explains difference in CO_2 between EOT and MMCO, despite similar $\delta^{18}{\rm O}$
- This entails a different construction of the $\delta^{18}{\rm O}$ signal between these episodes
- The result is contingent on how topographic changes are implemented in our ice sheet model

Correspondence to:

L. B. Stap, L.B.Stap@uu.nl

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The MMCO-EOT conundrum: Same benthic δ^{18} O, different CO₂

Lennert B. Stap¹, Roderik S. W. van de Wal¹, Bas De Boer², Richard Bintanja³, and Lucas J. Lourens⁴

¹Institute for Marine and Atmospheric Research Utrecht, Utrecht University, Utrecht, Netherlands, ²School of Earth and Environment, Faculty of Environment, University of Leeds, Leeds, UK, ³Royal Netherlands Meteorological Institute (KNMI), De Bilt, Netherlands, ⁴Department of Earth Sciences, Faculty of Geosciences, Utrecht University, Utrecht, Netherlands

Abstract Knowledge on climate change during the Cenozoic largely stems from benthic δ^{18} O records, which document combined effects of deep-sea temperature and ice volume. Information on CO₂ is expanding but remains uncertain and intermittent. Attempts to reconcile δ^{18} O, sea level, and CO₂ by studying proxy data suffer from paucity of data and apparent inconsistencies among different records. One outstanding issue is the difference suggested by proxy CO₂ data between the Eocene-Oligocene boundary (EOT) and the Middle-Miocene Climatic Optimum (MMCO), while similar levels of δ^{18} O are shown during these times. This conundrum implies changing relations between δ^{18} O, CO₂, and temperature over time. Here we use a coupled climate-ice sheet model, forced by two different benthic δ^{18} O records, to obtain continuous and mutually consistent records of δ^{18} O, CO₂, temperature, and sea level over the period 38 to 10 Myr ago. We show that the different CO₂ levels between the EOT and MMCO can be explained neither by the standard configuration of our model nor by altering the uncertain ablation parametrization on the East Antarctic Ice Sheet. However, we offer an explanation for the MMCO-EOT conundrum by considering erosion and/or tectonic movement of Antarctica, letting the topography evolve over time. A decreasing height of the Antarctic continent leads to higher surface temperatures, reducing the CO₂ needed to maintain the same ice volume. This also leads to an increasing contribution of ice volume to the δ^{18} O signal. This result is, however, dependent on how the topographic changes are implemented in our ice sheet model.

1. Introduction

Long-term climate change on Earth is documented by marine benthic δ^{18} O records [*Lisiecki and Raymo*, 2005; *Zachos et al.*, 2008; *Cramer et al.*, 2009]. They suggest a gradually cooling climate during the Cenozoic, overprinted by more rapid orbital-scale variability and transient excursions. However, the interpretation of the δ^{18} O signal is complicated by the fact that it comprises contributions from both deep-sea temperature and global ice volume [e.g., *Chappell and Shackleton*, 1986]. To disentangle the δ^{18} O signal, independent records of either deep-sea temperature [*Lear et al.*, 2000; *Sosdian and Rosenthal*, 2009; *Elderfield et al.*, 2012] or sea level [*Miller et al.*, 2005a; *Kominz et al.*, 2008; *Rohling et al.*, 2014] are required, which are still only scarcely available and possess large uncertainties.

The forcing of the documented climate change is externally determined by variations in solar insolation [*Milankovitch*, 1930]. However, these are not strong enough to explain the amplitude of observed climate variability without any further climate feedbacks [*Imbrie et al.*, 1992]. Furthermore, in contrast to δ^{18} O, the spectral power of solar insolation does not significantly change during the Cenozoic [*Zachos et al.*, 2008; *Laskar et al.*, 2004]. Therefore, insolation variability can neither explain the gradual cooling, nor the excursions documented by benthic δ^{18} O records. This indicates that internal variability and feedbacks play a crucial role in climate change over time. One factor concerns concentrations of greenhouse gases, such as CO₂, which heat up the atmosphere by trapping long-wave radiation. While a high-fidelity CO₂ record over the past 800 kyr is derived from ice cores [*EPICA community members*, 2006], proxy data records for CO₂ before that time are of low resolution, intermittent, and mutually disagreeing [*Beerling and Royer*, 2011; *Köhler et al.*, 2015]. Attempts to reconcile sea level, temperature, and CO₂ proxy data are hampered by this paucity of data and the mutual inconsistency of data records [*Foster and Rohling*, 2013; *Gasson et al.*, 2012].

©2016. American Geophysical Union. All Rights Reserved. An alternative approach to these issues, involving the use of models, is undertaken in this study. Models have the advantage of providing mutually consistent and continuous records. In earlier studies, the benthic

 $δ^{18}$ O signal was deconvoluted by using an inverse routine to force a one-dimensional ice sheet model. In this manner, sea level and Northern Hemispheric temperature simulations were obtained over the past 40 Myr [*De Boer et al.*, 2010]. On shorter time scales, this methodology was also applied to a three-dimensional ice sheet model with realistic topography [*Bintanja et al.*, 2005; *Bintanja and Van de Wal*, 2008; *De Boer et al.*, 2013, 2014]. A limitation of these stand-alone ice sheet model studies was that information on climate only consisted of globally uniform perturbations to preindustrial temperatures. Moreover, CO₂ was not explicitly incorporated in the model. Therefore, a hybrid model-data approach was taken by *Van de Wal et al.* [2011] to continuously reconstruct a consistent CO₂ record over the past 20 Myr. They assembled various proxy-CO₂ data records and inferred a log linear relation between Northern Hemispheric temperature and CO₂. In a subsequent study, the one-dimensional ice sheet model was coupled to a zonally averaged energy balance climate model [*Stap et al.*, 2014]. The inverse benthic $δ^{18}$ O routine was modified to simulate CO₂ over the past 5 Myr [*Stap et al.*, 2016], using the stacked benthic δ^{18} O record of *Lisiecki and Raymo* [2005] as input. They simulated benthic δ^{18} O, sea level, temperature, and CO₂ for the first time in a fully coupled fashion over a multimillion year time span.

In this study, we apply this method of deconvoluting δ^{18} O, inversely simulating CO₂ and thereby reconciling these variables, to the period 38 Myr ago to 10 Myr ago. Over this time, we force the same model as *Stap et al.* [2016], explained in section 2, using the stacked benthic δ^{18} O records of *Zachos et al.* [2008] and *Cramer et al.* [2009]. Specifically, we focus on two important warm spells: the time right before the Eocene-Oligocene Transition (EOT; 33.9 Myr ago) and the Middle-Miocene Climatic Optimum (MMCO; 17 to 15 Myr ago). Both these episodes encompass similarly low benthic δ^{18} O values, suggesting comparable warm climates. However, proxy-CO₂ data for right before the EOT [*Pagani et al.*, 2011; *Pearson et al.*, 2009; *Retallack*, 2009] show significantly higher values than for the MMCO [*Pagani et al.*, 1999; *Foster et al.*, 2012; *Greenop et al.*, 2014; *Kürschner et al.*, 2008]. Although earlier modeling studies have addressed the EOT [e.g., *DeConto and Pollard*, 2003; *Pollard and DeConto*, 2005; *Wilson et al.*, 2013; *Ladant et al.*, 2014; *Gasson et al.*, 2014] as well as the MMCO [e.g., *You et al.*, 2009; *Langebroek et al.*, 2009, 2010; *Henrot et al.*, 2010; *Goldner et al.*, 2014], we are the first to transiently simulate a time span covering both events. In doing so, we obtain a consistent representation of the climate during these important times.

We investigate whether we can explain the MMCO-EOT conundrum of similar benthic δ^{18} O but different CO₂, for instance, through a change in the contributions of ice volume and deep-sea temperature to the benthic δ^{18} O signal. In section 3.1, where we use the standard settings of the model and a sensitivity experiment, we do not simulate a difference in CO₂ between right before the EOT and the MMCO. In section 3.2 we analyze simulations with Antarctic topography varying over time, to address the influence of erosion and tectonics. This offers a plausible explanation for the conundrum. However, this result is contingent on the way 3-D topographic changes are transferred to our simplified 1-D model. In section 3.3, we discuss some further factors that can affect the relation between benthic δ^{18} O, which are not incorporated in our setup. Conclusions are presented in section 4.

2. Model and Methods

We use a coupled climate-ice sheet model [*Stap et al.*, 2014], incorporating a routine to calculate benthic δ^{18} O (Figure 1). This model consists of a zonally averaged energy balance climate model [*North*, 1975; *Bintanja*, 1997] coupled to one-dimensional models of the five major Cenozoic ice sheets (Eurasian, North American, Greenland, West Antarctic, and East Antarctic ice sheets) [*De Boer et al.*, 2010]. The climate model has a 5° latitudinal resolution and includes an ocean component with six vertical layers, which simulates the large-scale meridional ocean overturning and contains a 1.25° resolution thermodynamical sea ice routine. The ocean overturning strength varies as a function of the polar to equatorial sea surface temperature difference [*Stap et al.*, 2014]. The calculated local atmospheric surface temperatures are provided to the surface mass balance routine of the ice sheet models. Herein, precipitation (*P*) follows from present-day precipitation (*P*₀), corrected for surface air temperature (*T*), and ice sheet radius (*R*) with respect to a critical radius (*R*_c):

$$P = P_0 e^{0.04T - R/R_c}.$$
 (1)

Accumulation is calculated as a temperature-dependent fraction of precipitation. Ablation *M* is obtained using an insolation-temperature melt equation [e.g., *Van den Berg et al.*, 2008]:

$$M = [10T + 0.513(1 - \alpha)Q + C_{abl}]/100,$$
⁽²⁾



Figure 1. Schematic overview of the coupled model. Information on local temperature (*T*) is transferred from the climate model to the ice sheet model every 500 model years. The ice sheet model returns surface height and ice volume changes to the climate model. Also, every 500 years, the inverse routine is used to obtain the CO_2 input for the climate model from the difference between modeled benthic $\delta^{18}O$ and an observed $\delta^{18}O$ value 500 years later. For the benthic $\delta^{18}O$ observations the stacked records of *Zachos et al.* [2008, Z08] and *Cramer et al.* [2009, C09] are used. Insolation input [*Laskar et al.*, 2004] is updated every 1000 years.

where T, α , and Q are local temperature, surface albedo, and radiation, incorporating Milankovitch radiation variations, respectively. Parameter C_{abl} determines the ablation threshold. For the East Antarctic ice sheet, it is set to -30 in the reference simulation; in the simulations with increased ablation it is increased to -5. The ice sheet model calculates ice flow using the Shallow Ice Approximation (SIA). The resulting surface height and surface-type changes are forwarded to the climate model, effectuating the surface-height-temperature and albedo-temperature feedbacks. Exchange of variables between the climate and ice sheet models takes place every 500 model years.

The coupled model is forced with insolation data [*Laskar et al.*, 2004], which is updated every 1000 years. In forward mode, forced by CO_2 prescribed by the EPICA Dome C record [*EPICA community members*, 2004], the coupled model is capable of reproducing the glacial-interglacial cycles of the past 800 kyr accurately in terms of temperature and sea level [*Stap et al.*, 2014]. In inverse mode, as used in this study, the CO_2 concentrations are obtained every 500 years from

$$CO_2 = \overline{CO_2} \times \exp\left[c \times \left\{\delta^{18}O(t) - \delta^{18}O_{obs}(t+0.5kyr)\right\}\right],$$
(3)

where $\delta^{18}O(t)$ is the benthic $\delta^{18}O$ calculated by the model as a function of ice temperatures, ice thicknesses, and deep-sea temperature [*De Boer et al.*, 2010] and $\delta^{18}O_{obs}(t + 0.5kyr)$ is the observed $\delta^{18}O$ value 500 years later. Furthermore, $\overline{CO_2}$ is the mean CO_2 concentration of the preceding 15 kyr, and *c* a strength-determining parameter [*Stap et al.*, 2016]. In earlier work, this inverse routine was shown to reconstruct CO_2 over the past 800 kyr in good agreement with the EPICA Dome C ice core record [*EPICA community members*, 2004; *Stap et al.*, 2016]. Figure 1 shows the coupling between the models, and the forcing, schematically.

In this study, we use the stacked records of Zachos et al. [2008, Z08] and Cramer et al. [2009, C09] for the observed benthic δ^{18} O record. These stacked records possess uncertainty due to the fact that they are composed of data records originating from different locations, oceanic settings, and foraminiferal species [Zachos et al., 2008; Cramer et al., 2009]. Nonetheless, we assume them to represent global average conditions. We simulate continuous mutually consistent records of benthic δ^{18} O, CO₂, sea level (ice volume equivalent), and global temperature over the period 38 to 10 Myr ago. In this way, we obtain simulations that transiently cover both the EOT and the MMCO, which are the focus of this study. On the original 1000 year resolution (as used in *Stap et al.* [2016]), the reference run shows larger variability in δ^{18} O than Z08 (RMSE = 0.42‰). This is because the inverse method yields too much variability in CO₂, leading to oversized fluctuations in δ^{18} O. However, if 40 kyr running averages are considered, the inverse method is very accurate ($r^2 = 0.99$, RMSE = 0.05‰).

			EAIS	Initial Topo	Topo Change
Name	Description	Forcing	C_{abl}	Change (m)	Period (Myr Ago)
REF	Reference run	Z08	-30	0	NA
CRAM	Different forcing	C09	-30	0	NA
ABL+	Increased ablation EAIS	Z08	-5	0	NA
TOPO-	Higher Antarctica decreased	Z08	-5	520	34-0
TOPO+	Higher Antarctica increased	Z08	-5	790	34-0
TOPO-Q	Higher Antarctica accelerated	Z08	-5	655	34–20

 Table 1. Description of the Transient Model Runs Over the Period 38 to 10 Myr Ago^a

^aForcing records are the stacked benthic δ^{18} O records of Zachos et al. [2008, Z08] and Cramer et al. [2009, C09]. The ablation threshold parameter (C_{abl}) is defined in equation (2). Topographic changes are applied to the area 60–90°S in the climate model and to the ice sheet models for the East and West Antarctic Ice Sheets (EAIS and WAIS). NA = not applicable.

Therefore, all model input and output data are displayed as 40 kyr running averages in this study, unless otherwise stated. Hence, variations on smaller time scales than 40 kyr can and will not be assessed. Table 1 summarizes all the model runs we performed.

3. Results and Discussion

3.1. The MMCO-EOT Conundrum

Our reference simulation is obtained by running the model over the period 38 to 10 Myr ago, after a spin-up of 2 Myr (Table 1). The stacked benthic δ^{18} O record of Zachos et al. [2008, Z08] is used as input. This simulation shows an only partly glaciated Antarctica, sustained by high CO₂ concentrations of 650 to 750 ppm, until the Eocene-Oligocene Transition (EOT; 33.9 Myr ago) (Figure 2). At this time, a CO₂ drop to 325 ppm leads to full-scale glaciation of Antarctica and a 2.6 K global temperature cooling. After this event, the East Antarctic ice sheet remains very variable in size (sea level varying between 0 and 50 m sea level equivalent (s.l.e.) above preindustrial (PI) level) during the Oligocene and early Miocene. This is sustained by CO₂ levels between 400 and 750 ppm, with global mean temperature varying between 289 and 292 K. At the Middle-Miocene Climatic Optimum (MMCO; 17 Myr ago), sea level and CO₂ peak at 55 m above PI and 750 ppm, respectively. Thus, in this simulation the CO₂ concentrations at the MMCO and right before the EOT are similar, meaning there is no sign of hysteresis in the relation between δ^{18} O and CO₂. While this is in agreement with the findings of Hansen et al. [2013], who imposed a nonchanging relation between CO₂ and benthic δ^{18} O via temperature, proxy-CO₂ data suggest different CO₂ levels between these two episodes. This is visible in Figure 3 (red line), where we focus on two periods that enclose these warm spells: (a) middle Miocene: 18 to 13 Myr ago and (b) early Oligocene: 35 to 30 Myr ago. We compare our simulated CO₂ to proxy data based on alkenones [Pagani et al., 1999, 2011], boron isotopes [Foster et al., 2012; Greenop et al., 2014; Pearson et al., 2009], and stomata [Retallack, 2009; Kürschner et al., 2008]. Before the EOT, there is broad agreement on CO₂ levels ranging from 800 to 1000 ppm between the model results and the data based on alkenones [Pagani et al., 1999] and stomata [Retallack, 2009] (Figure 3b). The boron isotope-based reconstruction of Pearson et al. [2009] shows slightly higher CO₂ values. Our modeled glaciation threshold CO₂ concentration of approximately 700 ppm is within the 560 to 920 ppm range found by Gasson et al. [2014], who studied different combinations of an ice dynamical model coupled to seven climate models (Figure 3b, yellow shading). During the early Oligocene period after the EOT, our reconstructed CO_2 is generally lower than proxy data. The paradoxical simultaneous occurrence of relatively low CO₂ and low δ^{18} O, indicative of high sea level, during the MMCO is very hard to capture. Earlier modeling studies either showed high CO₂ and high sea level [Langebroek et al., 2009, 2010; Henrot et al., 2010; Goldner et al., 2014] or low CO₂ and relatively low sea level [You et al., 2009]. In our reference run, we do not simulate low sea level, because the δ^{18} O in the forcing record [Zachos et al., 2008] is too low. Our simulated sea level is in fact similarly high as before the EOT, implying a similar contribution of sea level and temperature to the δ^{18} O signal. In contrast to our simulation, proxy data show much lower CO₂ concentrations during the MMCO than before the EOT (Figure 3a), although the uncertainties in these reconstructions are substantial. Stomata-based data [Kürschner et al., 2008] ranges between 293 and 469 ppm, while alkenone-based [Pagani et al., 1999] and boron-isotope-based [Foster et al., 2012; Greenop et al., 2014] data show generally lower values of 195 to 540 ppm during the MMCO. The question now arises, what could cause this disagreement between model and data?



Figure 2. Time series of benthic δ^{18} O, CO₂, sea level, and global temperature. (a) Simulated benthic δ^{18} O, (b) simulated CO₂, (c) simulated sea level in meters above present day, and (d) simulated global mean temperature (T_{glob}). Shown are 40 kyr running averages of the reference run (REF). Dotted lines represent preindustrial (PI) values. Highlighted in yellow are the early Oligocene period (35 to 30 Myr ago) and the middle Miocene period (18 to 13 Myr ago).

Obviously, our reconstruction depends strongly on the δ^{18} O record that is used as forcing. In another run (CRAM), we force the coupled model with the stacked δ^{18} O record of *Cramer et al.* [2009, C09] (Table 1). In contrast to Z08, this compilation takes into account differences between ocean basins in the stacking method. The main difference with REF in simulated CO₂ is the timing of full-scale Antarctic glaciation at the EOT and after the MMCO; in simulation CRAM this takes place earlier than in REF. Although the higher δ^{18} O during the middle to late Miocene in C09 is reflected in lower sea level (not shown), the difference in CO₂ between REF and CRAM is relatively small. This is because the Antarctic Ice Sheet is very sensitive in this CO₂ regime; a small shift in CO₂ leads to a large ice volume change. Hence, both by forcing with C09 and Z08, the simulated difference between CO₂ levels before the EOT and during the MMCO are qualitatively similar.

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Figure 3. Simulated CO₂ concentrations. Simulated CO₂ for REF (red), CRAM (cyan), and ABL+ (black) for (a) the middle Miocene and (b) the early Oligocene. Symbols indicate proxy-data reconstructions based on alkenones [*Pagani et al.*, 1999, 2011] (orange asterisks), boron isotopes [*Foster et al.*, 2012; *Greenop et al.*, 2014; *Pearson et al.*, 2009] (blue pluses), and stomata data [*Retallack*, 2009; *Kürschner et al.*, 2008] (green crosses). Yellow shading indicates the range of Antarctic glaciation values found by *Gasson et al.* [2014].

An important model uncertainty resides in the mass balance parametrization of the East Antarctic Ice Sheet (EAIS). *Stap et al.* [2016] already showed that increasing the ablation by implementing a lower ablation threshold, leads to a more dynamic EAIS during the Pliocene. We perform a similar experiment (ABL+) forcing the model over the period 38 to 10 Myr ago using Z08 and lowering (in absolute sense) the ablation threshold parameter from its reference value -30 to -5 (Table 1). The increased ablation in this run leads to a decreased threshold CO₂ values for East Antarctic glaciation and deglaciation. Consequently, the resulting CO₂ levels right before the EOT and during the MMCO are lower in the ABL+ simulation than in the REF simulation. This also induces a new balance in the contributions to the benthic δ^{18} O signal. Since CO₂ is lower, deep-sea temperatures are lower, and hence, ice volume must contribute less to the signal and simulated sea level is somewhat higher (not shown). The simulated lower MMCO CO₂ levels are in better agreement with proxy data (Figure 3a; black line). Contrarily, late Eocene and early Oligocene CO₂ concentrations are now unrealistically low (Figure 3b; black line). In other words, a difference in CO₂ levels between both periods is still not simulated, so the conundrum remains.

The similarity between the early Oligocene and middle Miocene periods in this run is also reflected by the polar amplification in the Southern Hemisphere (Figure 4a). In agreement with data reconstructions [e.g., *Goldner et al.*, 2014; *Bijl et al.*, 2009], we simulate reduced equator-to-South-Pole temperature gradients



Figure 4. Transient polar amplification. Relation between global and Antarctic (60° to 90°S) mean annual surface temperature, during the middle Miocene (blue dots) and the early Oligocene (red dots). For (a) ABL+ and (b) TOPO-Q.



Figure 5. Polar amplification. Relation between global and Antarctic (60 to 90°S) mean annual surface temperature in schematic equilibrium runs using the standard configuration (red), PD land ice conditions (blue), PD ocean overturning strength (black), and PD land ice and ocean overturning strength (orange).

right before the EOT and during the MMCO. This is demonstrated by more strongly increased Antarctic temperatures (up to 9 K) compared to the global mean (up to 3 K) with respect to PD. To investigate the factors determining polar amplification, we additionally perform schematic equilibrium runs of the model. In these runs, we first increase CO_2 in steps of 50 ppm per 50 kyr from 450 ppm to 1200 ppm and then oppositely decrease the concentration back to 450 ppm. Insolation is kept at presentday configuration throughout all these schematic runs. If we do not take into account changing land ice and changing ocean overturning strength, the polar amplification is relatively small and uniform (Figure 5, orange line). Allowing for ocean overturning strength variations increases the amplitude of polar amplification as is indicated in Figure 5 by a slightly steeper blue line. Switching on land ice changes also leads to nonuniform polar amplification (Figure 5, black

line). In the latter case, it becomes stronger in the temperature regime where the Antarctic Ice Sheet is dynamic and hence very sensitive to temperature changes. In addition, significant hysteresis is introduced by the surface-height-temperature and albedo-temperature feedbacks, as is shown by the difference between the upward and the downward branches. The separate effects of ocean overturning strength and land ice changes combine more or less linearly (Figure 5, red line).

3.2. Influence of Changing Antarctic Topography

3.2.1. Bulk Effect

In all the runs discussed so far we use present-day topography as boundary condition by assuming a local isostatic equilibrium. However, as a consequence of, e.g., tectonics and erosion, ice-free topography has significantly changed since the late Eocene [Rugenstein et al., 2014, and references therein]. Particularly, the Antarctic topographical changes may have had a significant influence on the dynamics of the Antarctic ice sheet [Siegert, 2008; Oerlemans, 1984; Huybrechts, 1993; Wilson et al., 2013; Gasson et al., 2015; Austermann et al., 2015; Gasson et al., 2016]. Wilson et al. [2012] calculated late Eocene Antarctic topography relative to the isostatically rebounded BEDMAP present-day topography [Lythe et al., 2001], using a set of tools including ice sheet-erosion models, models of thermal subsidence, and plate movement. They constructed a minimum and maximum estimated late Eocene topography difference from present day. This study shows eroded volumes that lower the Antarctic continent by 520 to 790 m on average. We perform experiments, where we increase the initial height of the Antarctic continent 38 Myr ago by 520 m (TOPO–) and by 790 m (TOPO+) in the climate model (60 to 90°S) as well as in the ice sheet models of East and West Antarctica. We thereby ignore isostatic adjustment of the eroded material, which counteract lowering of the surface. Our modeling effort therefore probably overestimates the effect of erosion and should be seen as a proof of concept rather than a rigorous quantification. Since not much is known of Antarctic topography evolution over time, we impose a linear decrease of the height difference over time. Data-based evidence suggests that erosion on Antarctica was very weak until glaciation of the continent [Thomson et al., 2013]. Therefore, we start this decrease at 34 Myr ago. The rate of height decrease is such that zero difference would be obtained at present day. However, because Antarctica may have eroded more strongly during the early parts of the Cenozoic [Tochilin et al., 2012], we perform a separate run (TOPO-Q) where the rate of topographic change is quicker. In that run, we prescribe a height decrease of 655 m (the average of 520 and 790 m) linearly over the shorter period 34 to 20 Myr ago. In all these runs, we use the enhanced ablation parameterization of run ABL+ (Table 1), since this gives the best agreement with proxy data during the MMCO.



Figure 6. Simulated CO_2 concentrations. Simulated CO_2 for REF (red), TOPO-Q (green), and the range bounded by TOPO- and TOPO+ (blue shading) for (a) the middle Miocene and (b) the early Oligocene. Yellow shading indicates the range of Antarctic glaciation values found by *Gasson et al.* [2014]. Symbols as in Figure 3.

A higher Antarctic surface at the EOT generally experiences lower surface air temperatures at the same CO₂ level. Therefore, the threshold CO₂ value for (de)glaciation is higher. Because the model is forced by the same values of benthic δ^{18} O, in the TOPO experiments we obtain compensating higher levels of CO₂ than in the ABL+ experiment (Figure 6b); even higher than simulated in the REF experiment. The modeled CO₂ ranges from 900 to 1400 ppm before and from 500 to 900 ppm after the EOT, in better agreement with proxy data. At the MMCO, the topographic difference between the TOPO runs and ABL+ has reduced, and simulated CO₂ levels are closer together (Figure 6a). The TOPO– and TOPO+ runs show CO₂ levels almost equal to REF (Figure 6a; blue shading). The CO₂ range spanned by TOPO– and TOPO+ is also smaller because the topographic difference between these runs is reduced. The CO₂ levels simulated in run TOPO-Q are similarly high as ABL+ at the MMCO, since there is no difference in topography with ABL+ at this time anymore (Figure 6a; green line). Any difference in CO₂ is therefore explained by transient effects. Hence, a changing Antarctic surface topography leads to a difference in CO₂ between right before the EOT and during the MMCO, while similar benthic δ^{18} O is simulated.

A changing relation between Antarctic and global mean temperatures over time is also shown by Figure 4b. During the early Oligocene period, much higher global mean temperatures are simulated by the TOPO-Q simulation than during the middle Miocene period. While similar Antarctic surface temperatures are needed to glaciate the continent, the higher surface elevation ensures that Antarctica is colder with respect to the rest of the globe. Therefore, Antarctica glaciates at higher global mean temperatures, and hence higher concentrations of CO_2 . Over the course of time, when the Antarctic surface elevation drops (Figure 7e), the CO_2 levels needed (Figure 7b) to simulate the same benthic $\delta^{18}O$ (Figure 7a) gradually decrease. This concurs with slowly declining global temperatures (Figure 7d). Additionally, simulated sea level before the EOT is lower than during the MMCO (Figure 7c), ~10 m lower than the 40–50 m above present day reconstructed by *Miller et al.* [2005b]. This is because at the EOT, CO_2 levels are increased, leading to warmer deep-sea temperatures. This ensures that a smaller part of the benthic $\delta^{18}O$ decrease with respect to the preindustrial value is accounted for by loss of land ice.

3.2.2. Separate Effect of Several Factors

Having shown the bulk effect of changing Antarctic topography, we recognize that we do not include local effects which may be of importance [*Gasson et al.*, 2015; *Jamieson et al.*, 2010]. The crude resolution of our climate model, as well as the idealized topography of our ice sheet model, hinders doing this. However, in the following we attempt to quantify the uncertainty pertaining to not including local effects by investigating three potentially important factors: (1) the effect of separate EAIS and WAIS height increases (separated E/W in Figure 8), (2) the effect of eliminating the land surface area increase caused by the height increase (decreased area in Figure 8), and (3) the effect of not changing the central height but instead reducing the slope of the EAIS and WAIS (reduced H_{cnt} in Figure 8). To study these effects, we perform three further tests over the shorter period 38 to 30 Myr ago. In these experiments, we do not include any transient changes. Therefore, we also perform a new standard run for comparison, where this time we increase the Antarctic topography by 655 m over the entire run (standard in Figure 8).



Figure 7. Time series of benthic δ^{18} O, CO₂, Antarctic surface elevation deviation from PD, and global temperature. (a) Simulated benthic δ^{18} O from TOPO-Q, (b) simulated CO₂ from TOPO-Q (red line) and from TOPO- and TOPO+ (orange shading), (c) simulated sea level in meters above present day from TOPO-Q (blue line) and from TOPO- and TOPO+ (cyan shading), (d) simulated global mean temperature (T_{glob}) from TOPO-Q (black line) and from TOPOand TOPO+ (grey shading), and (e) Antarctic surface elevation deviation from PD and from TOPO-Q (solid line) and TOPO- and TOPO+ (dashed lines). Shown are 40 kyr running averages. Dotted lines represent preindustrial (PI) values. Highlighted in yellow are the early Oligocene period (35 to 30 Myr ago) and the middle Miocene period (18 to 13 Myr ago).

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Figure 8. Simulated CO_2 concentrations and sea level. (a) Simulated CO_2 and (b) sea level for the runs performed in section 3.2.2. In red, the standard run with and increased Antarctic height of 655 m. In blue, the run with separate height increases of 401 m for the EAIS and 865 m for the WAIS. In black, the run with increased height of 655 m, but reduced surface area. In green, the run without central height (H_{CNT}) increase, but with a reduced slopes for the EAIS and WAIS.

In the first test, we make a distinction between the topography of the EAIS and the WAIS, as evidence suggests that the WAIS eroded more than the EAIS. We apply a height increase of 401 m to the EAIS, and 865 m to the WAIS. This is based on separating the eroded volumes of both these areas in Table 3 of *Wilson et al.* [2012]. Doing this slightly hinders the growth of land ice, leading to 101 ppm lower CO₂ concentrations (Figure 8a, blue line), and 4.8 m higher sea level (Figure 8b, blue line) averaged over the whole simulated period 38 to 30 Myr ago.

The 655 m height increase we apply facilitates the Antarctic surface area above sea level becoming 243% larger when there is no land ice (a contribution of 70 m sea level), and 235% larger at present-day sea level. This leads to an easier build up of more land ice. To quantify this effect, we perform a second test where we artificially eliminate this surface area increase. In addition to applying the 655 m height increase of the topography, we raise the sea level by the same amount. In this way, the Antarctic continent still experiences lower temperatures but does not have a larger land surface area. At high sea levels, there is not much difference with the new standard run (Figure 8, black lines). However, toward colder conditions, land ice growth is capped at approximately present-day sea level. In the new standard run, the Antarctic Ice Sheet grows larger, and sea level drops to -14.4 m at the end of the EOT. This indicates an Antarctic Ice Sheet size of 84.4 m s.l.e., $\sim 20\%$ larger than present day, in agreement with the 83.4 to 89.5 m s.l.e. range found by *Wilson et al.* [2013]. A smaller Antarctic Ice Sheet leads to a reduced benthic δ^{18} O signal. This is compensated by lower deep-sea temperatures, and hence decreased CO₂. After the EOT CO₂ is on average 105 ppm and maximally 159 ppm lower than in the new standard run.

Finally, in our model the glacial inception threshold may be contingent on the height of the top of the continental cone, where glaciation starts. Translated to the real world, this implies a dependency on the height of the mountainous regions of the Antarctic continent. These regions, however, may not have eroded as much as the lower lying parts of Antarctica [*Rose et al.*, 2013]. Therefore, in the third test we do not increase the central height of the continents but reduce the slopes such that the volume above sea level remains the same as in the new standard run. This means a slope of -0.474 instead of -1.0 m/km for the EAIS, and a slope of -0.158 instead of -1.1 m/km for the WAIS. Clearly, this has a very large effect on the simulated CO₂ (Figure 8a, green line) and sea level (Figure 8b, green line). In this simulation CO₂ and sea level are hardly different from the ABL+ results (not shown), being, respectively, much lower and higher than in the new standard run.

3.3. Other Influencing Factors

Other important factors that can affect the glaciation threshold CO_2 are ocean dynamics and vegetation dynamics. In our model, we use a simple meridional overturning scheme for the ocean [*Bintanja*, 1997; *Stap et al.*, 2014]. The strength of the overturning is determined by the density difference between equatorial and polar waters. In reality, however, a change in the structure of the ocean basins can also have a significant effect on ocean dynamics. For instance, closure of the Drake Passage has a lowering effect on the Antarctic glaciation CO_2 threshold at the EOT [*Kenneth*, 1977; *DeConto and Pollard*, 2003; *Toggweiler and Bjornsson*, 2000] by increasing ocean heat transport toward the South Pole. Another data study argued that a deeper global thermocline and reduced low-latitude gradients in sea surface temperature during the Miocene may explain an altered relation between CO_2 and temperature [*LaRiviere et al.*, 2012]. Furthermore, we ignore vegetation dynamics. In our model, the Antarctic continent is covered by grassland when it is ice free. If the vegetation

consisted (partly) of forests during the EOT, the CO_2 threshold for Antarctic glaciation could be much lower [*Liakka et al.*, 2014]. These mechanisms would favor higher CO_2 at the MMCO than at the EOT opposite to what the data suggest. Therefore, we leave analysis of the impact of these factors to studies using more sophisticated models that contain more detailed ocean and vegetation physics.

4. Conclusions

We have presented continuous mutually consistent simulations of benthic δ^{18} O, CO₂, sea level, and global temperature, over the period 38 to 10 Myr ago (Figure 2). These are obtained using a coupled climate-ice sheet model, with an inverse routine to get CO₂ from a stacked benthic δ^{18} O record [*Zachos et al.*, 2008; *Cramer et al.*, 2009]. This approach based on simplified models provides the unique possibility of reconciling these variables on multimillion year time scales.

We have focused on two warm spells, the EOT and the MMCO, where proxy data suggest different CO₂ levels, while benthic δ^{18} O values are similar. We find that this conundrum cannot be explained using the standard configuration of our model. We simulate similar CO₂ levels right before the EOT as during the MMCO. The contributions of ice volume and deep-sea temperature to the benthic δ^{18} O signal during these two episodes are also comparable. Forcing the model by the benthic δ^{18} O record of *Cramer et al.* [2009], which shows higher values during the middle Miocene, we do obtain lower sea level. The simulated CO₂ levels, however, are only slightly reduced. Lowering the uncertain ablation threshold for the EAIS leads to decreasing levels of CO₂ at both events and therefore offers no solution to the conundrum.

In order to offer a solution to the conundrum, we investigate the first-order effect of evolving Antarctic topography. For this purpose, we run the model starting with a height increase of 520 to 790 m based on the eroded volumes in *Wilson et al.* [2012] applied to the Antarctic continent. We diminish this height increase either slowly (to reach zero at present day) or more quickly (to reach zero at 20 Myr ago). Through the lapse rate effect, a higher surface at the EOT experiences lower temperatures. This means that the EAIS glaciates and deglaciates at higher CO₂ levels. Therefore, low benthic δ^{18} O values are reached at these increased CO₂ levels and contain a comparatively larger contribution from warmer deep-sea temperatures. Especially when the topographic difference is reduced at a relatively fast pace, this offers an explanation for differences in CO₂ between the EOT and MMCO while benthic δ^{18} O values are similar. Additionally, the larger land mass effectuated by the height increase can sustain a 20% larger Antarctic ice sheet. When spatial heterogeneity is included by applying different height increases for the EAIS and WAIS separately, the CO₂ difference is somewhat (~100 ppm) reduced.

The effect of increased Antarctic surface elevation is, however, strongly contingent on how the 3-D topographic changes are implemented in the 1-D ice sheet model. The threshold for (de)glaciation in our model is determined by the maximum height of the continental cone. If instead of applying a uniform height increase, this peak height is kept constant and only the slope of the Antarctic continent is reduced, the simulated EOT CO_2 is much lower. Therefore, to better constrain the influence of erosion and tectonic movement on climate, a 3-D ice sheet model should be used. Ideally, this model would be interactively coupled to an erosion model. Inclusion of isostatic adjustment should further help to obtain more realistic results.

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